

# Exploring the Galactic Warp Through Asymmetries in the Kinematics of the Galactic Disk

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[arXiv:2010.10398](https://arxiv.org/abs/2010.10398)

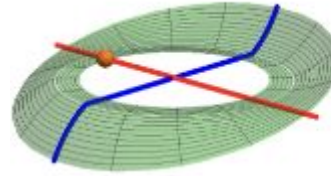
QR Code to Paper



**gaia**

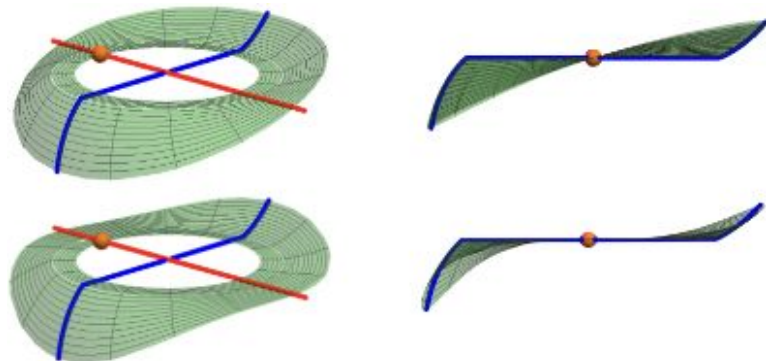


# Introduction



- Galactic Warp: the bending of Galactic disk
- Found in majority of spiral galaxies (Bosma 1978)
  - Long-live/repeatedly regenerated
- Limited to solar neighborhood before Gaia
- Cause: under debate
  - Interaction with satellite galaxies
    - Sgr dSph (Ibata & Razoumov 1998; Laporte et al. 2019)
    - LMC, SMC (Weinberg & Blitz 2006; Garavito-Camargo et al. 2019)
  - External torques of dark matter halos (Widrow et al. 2014)
  - Accretion of intergalactic matter (Ostriker & Binney 1989)
  - Misaligned dark matter halo (Sparke & Casertano 1988)
  - Intergalactic magnetic field (Battaner et al. 1990; Guijarro et al. 2010)

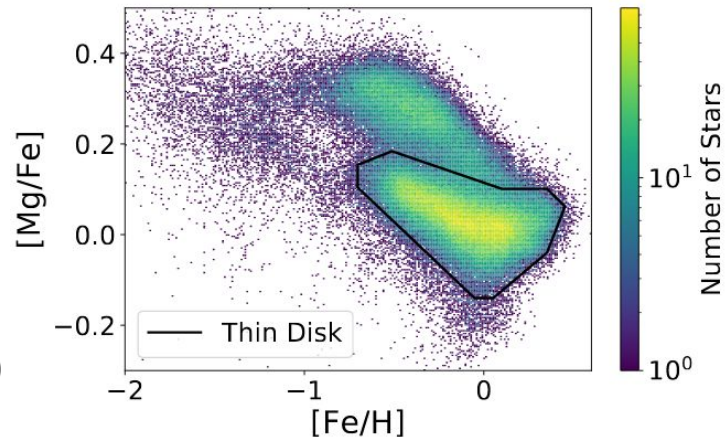
# Introduction



- Geometry: uncertain
  - Shape
    - Sine-lopsided or S-lopsided (Romero-Gomez et al. 2019)
  - Starting radius
    - Inside solar circle
      - J/K-band surface brightness fitting (Drimmel & Spergel 2001)
      - Gaia-TGAS (Schonrich & Dehnen 2018), LAMOST-TGAS (Huang et al. 2018)
    - Outside solar circle
      - Star count (Reyle et al. 2009; Derriere & Robin 2001)
      - Gaia DR2 kinematics (Romero-Gomez et al. 2019)
- This work
  - Kinematics from Gaia DR2
  - Chemistry from APOGEE
  - Explore asymmetries in the outer Galactic disk

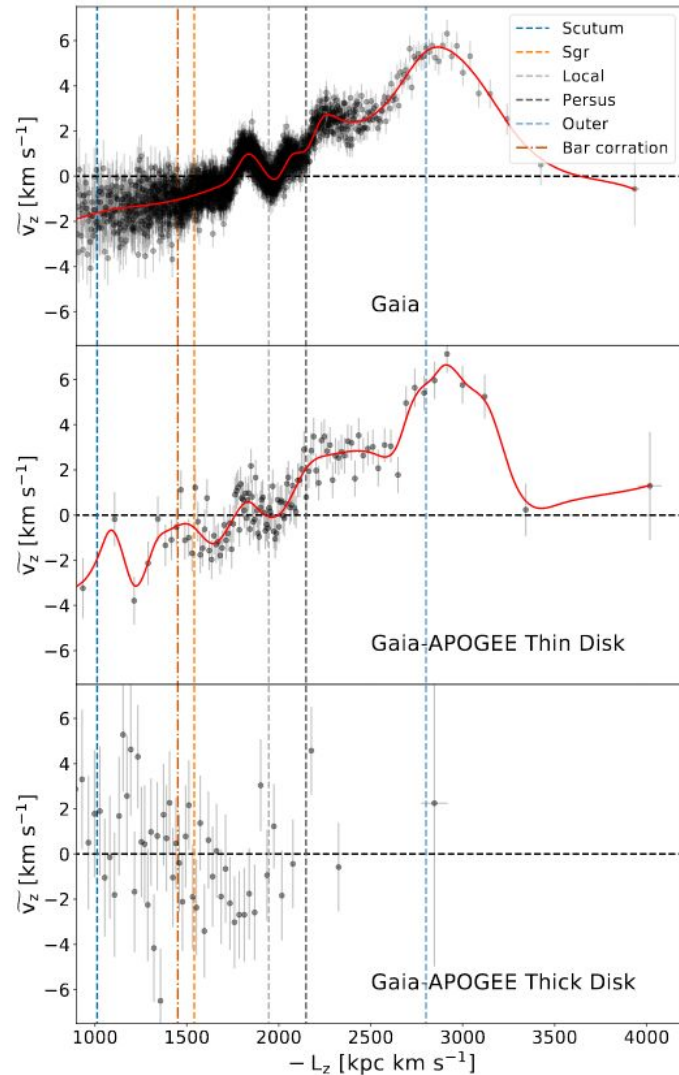
# Data

- 2 data-sets
- Gaia
  - Proper motion + StarHorse distance (Gaia) + RV (Gaia)
  - $G < 15$  mag
  - Removal of stars with suspect photometry and RV based on  $< 4$  visits
- Gaia-APOGEE
  - PM (Gaia) + StarHorse distance (Gaia) + RV (APOGEE) + chemistry information (APOGEE)
  - Chemically selected thin disk
- StarHorse distance
  - Bayesian estimation of stellar parameters, distances and extinctions with the combination of photometric and parallax information from Gaia, Pan-STARRS1, 2MASS, and AllWISE



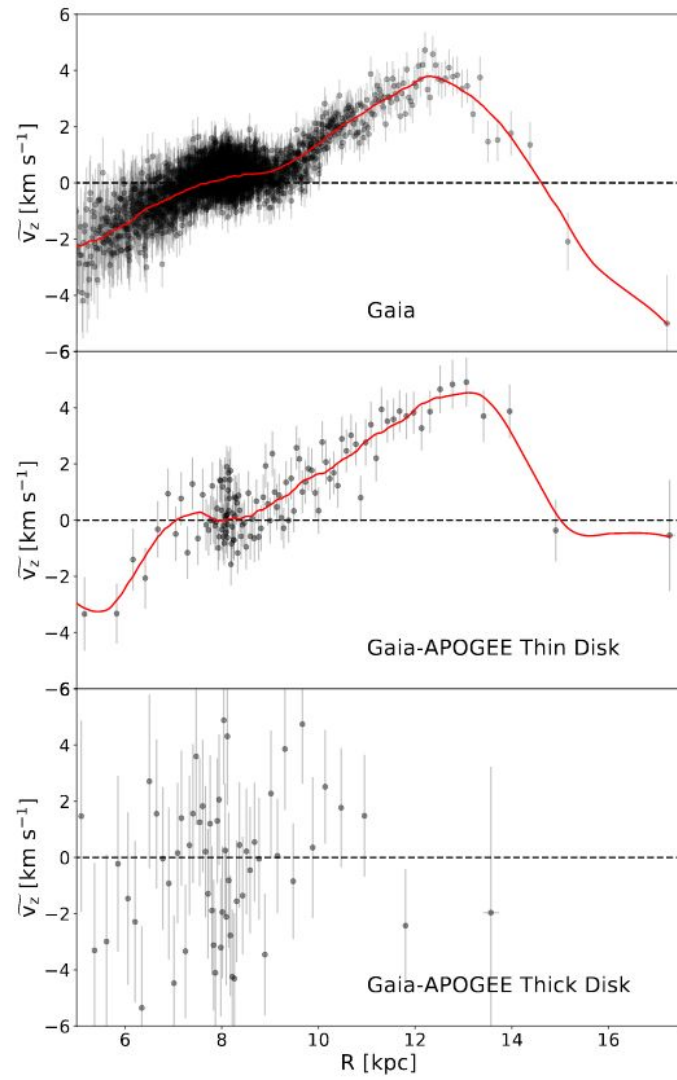
# Structures in Vertical Velocity

- $V_z$  increase with  $L_z$ 
  - $L_z < 2800$  kpc km/s
  - Previously observed and attributed to Galactic warp
- Peak  $V_z = 6$  km/s
  - $L_z \sim 2800$  kpc km/s
- Drops to 0
  - $L_z \sim 3500$  kpc km/s
- First time discovery of the drop-off
- Substructure (wiggles)
  - $L_z \sim 1700, 2000, 2200$  kpc km/s
  - Some aligned with spiral arms



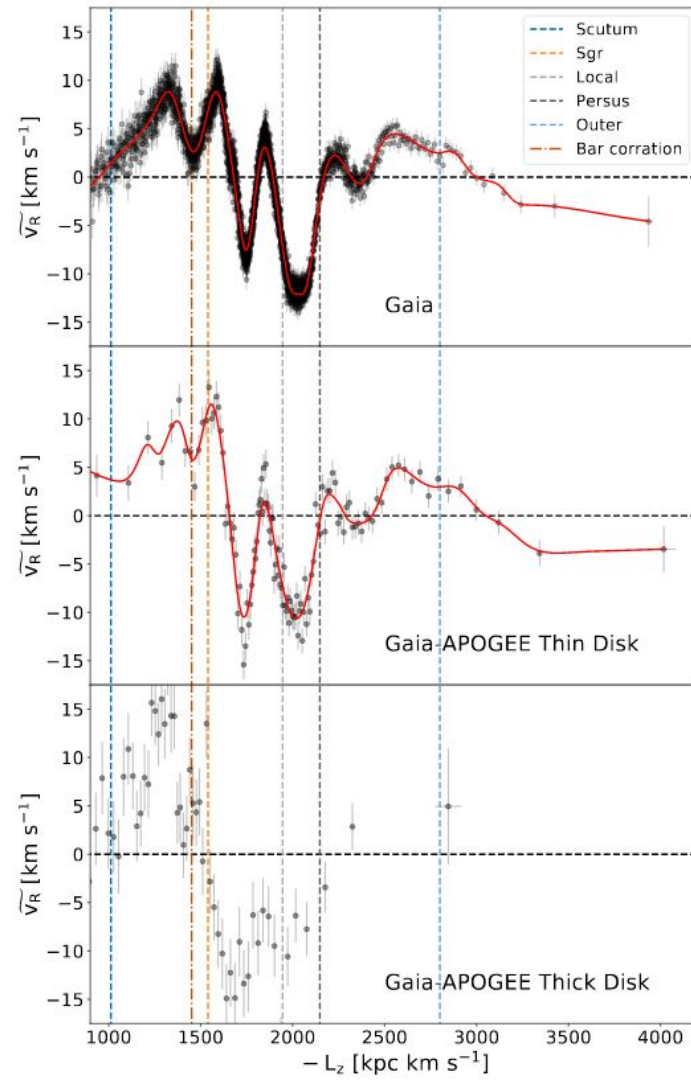
# Structures in Vertical Velocity

- General trend also exist when viewed against Galactocentric radius
- Substructures smeared out since  $R$  is not an integral of motion
- Plateau of  $V_z \sim 0$  at solar neighborhood



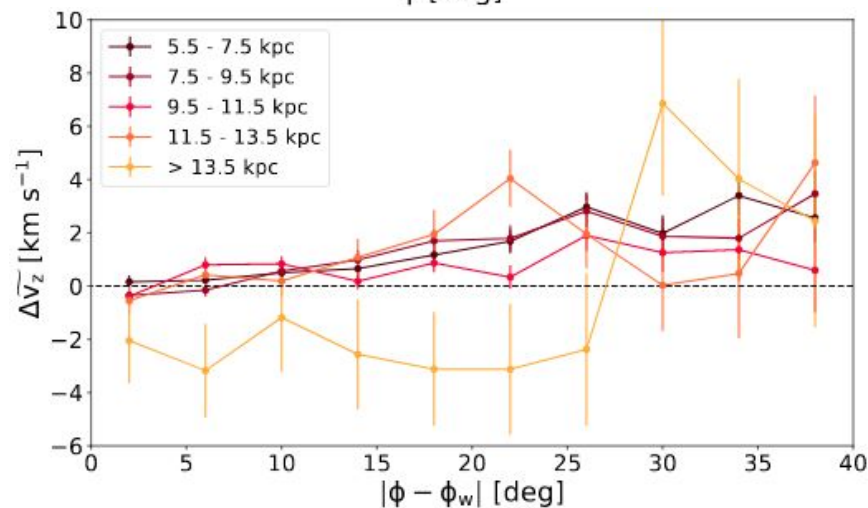
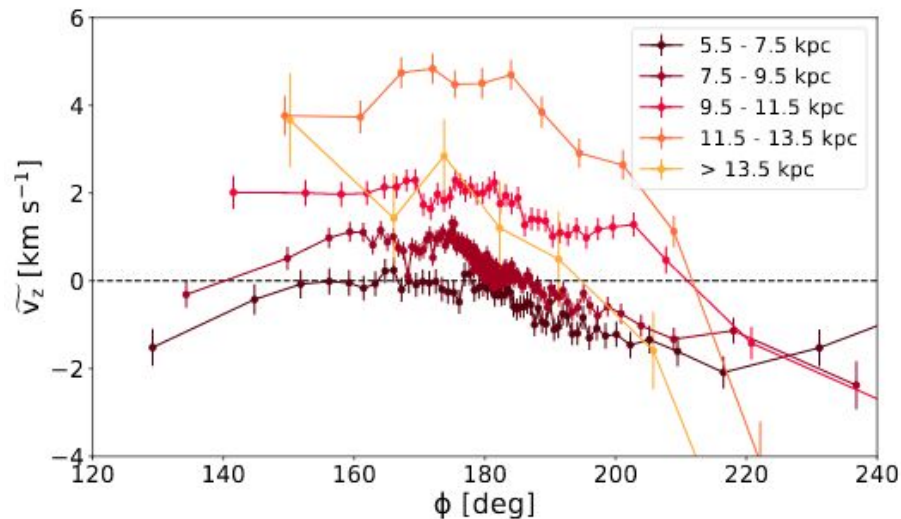
# Structures of Radial Velocity

- Large-amplitude ( $\sim 13$  km/s) ripples detected
- Phase difference with vertical velocity
- Not all peak/valley aligned with spiral arms



# Lopsided Warp

- Romero-Gomez et al. 2019: up and down sides are not symmetric -> lopsided
- Peak velocity at  $\sim 170$  deg
- Rate of change in vertical velocity not symmetric with respect to maximum point





# Modeling the Observed Vertical Kinematics

- Physical reason for the drop-off
- Stellar number density separable in R and z:  $n(R, z) = f(R)g(z)$
- Warp precessing in galactic-rotation direction:  $z_0 = h(R) \sin(\phi + \phi_w + \omega_p t)$

$$h(R) = \begin{cases} 0 & R \leq R_1 \\ \frac{R_w}{R_2 - R_1} (R - R_w)^\alpha & R_1 < R \leq R_2 \\ h(R_2) + \frac{dh}{dR} \Big|_{R=R_2-R_1} (R - R_2) & R > R_2 \end{cases}$$

- The first Jeans Equation:  $\frac{\partial n'}{\partial t} + \frac{\partial(n' \overline{v_R})}{\partial R} + \frac{1}{R} \frac{\partial(n' \overline{v_\phi})}{\partial \phi} + \frac{\partial(n' \overline{v_z})}{\partial z} = 0$
- Axis-symmetry  $\frac{\partial \overline{v_\phi}}{\partial \phi} = 0$  and  $\frac{\partial \overline{v_z}}{\partial z} = 0$  but a non-zero radial velocity

$$\begin{aligned} & \left[ \overline{v_z} - \left( \frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos(\phi + \phi_w + \omega_p t) - \right. \\ & \quad \left. \overline{v_R} \frac{dh}{dR} \sin(\phi + \phi_w + \omega_p t) \right] f(R) \frac{dg}{dz} \\ & + \overline{v_R} \frac{df}{dR} g(z - h(R) \sin(\phi + \phi_w + \omega_p t)) + \frac{\partial \overline{v_R}}{\partial R} n' = 0 \end{aligned}$$

# Modeling the Observed Vertical Kinematics

- If the Galaxy has the same IMF everywhere, the average mass of stars would also be a constant, thus the number density is proportional to mass density

$$n(R, z) = n_0 \exp\left(-\frac{|z|}{z_h} - \frac{R}{R_h}\right)$$

- Plug in and crunch:  $\frac{\partial \overline{v_R}}{\partial R} = \frac{\overline{v_R}}{R_h} + \frac{\text{sign}[z - z_0]}{z_h} \left[ \overline{v_z} - \left( \frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos \theta - \overline{v_R} \frac{dh}{dR} \sin \theta \right]$

where  $z_h$  is the scale height and  $R_h$  is the scale length

- Population is symmetric about  $z_0$ , so average  $Z > Z_0$  and  $Z < Z_0$

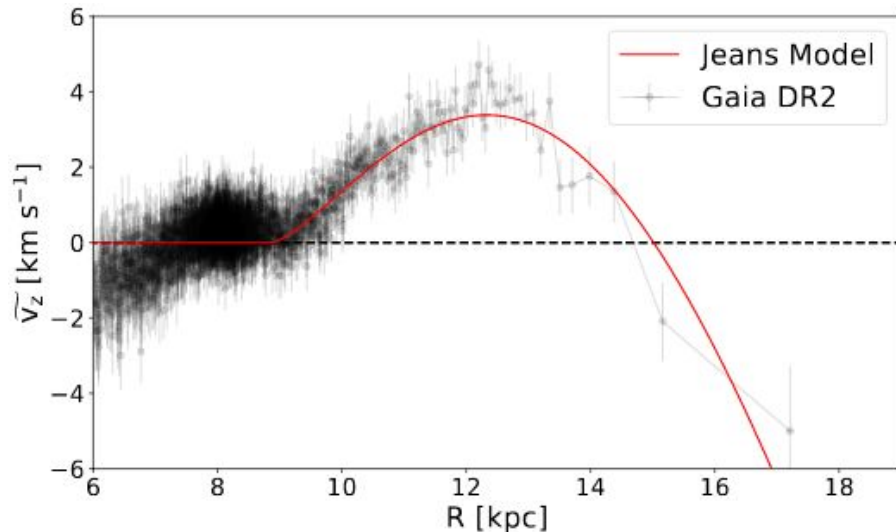
$$\overline{v_z} = \left( \frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos \theta + \overline{v_R} \frac{dh}{dR} \sin \theta$$

# Modeling the Observed Vertical Kinematics

- MCMC fitting

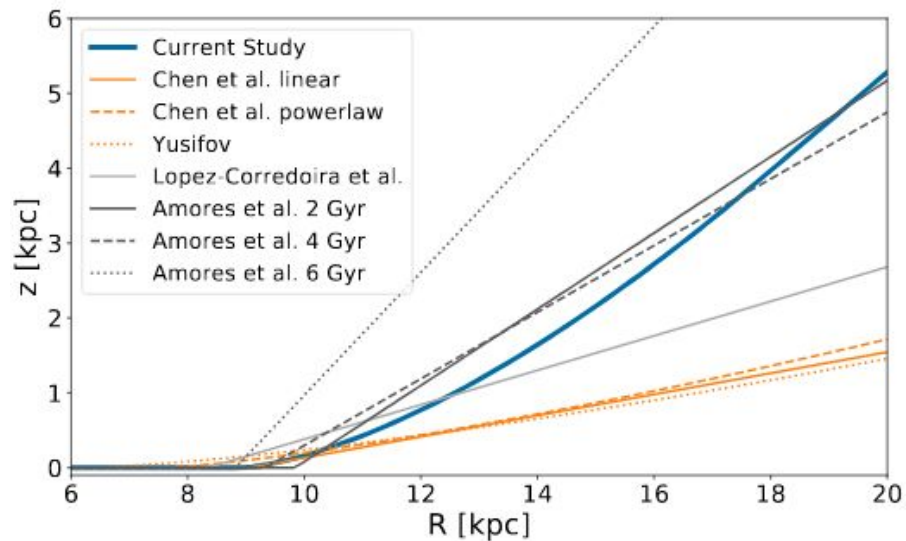
$$\left\{ \begin{array}{l} \omega_p = -13.57^{+0.20}_{-0.18} \text{ km s}^{-1} \text{ kpc}^{-1} \\ R_1 = 8.87^{+0.08}_{-0.09} \text{ kpc} \\ R_2 = 17.78^{+1.56}_{-1.86} \text{ kpc} \\ R_w = 1.20^{+0.28}_{-0.26} \text{ kpc}^{2-\alpha} \\ \alpha = 1.53^{+0.10}_{-0.09} \end{array} \right.$$

- Consistent with Poggio et al. 2020
- Not sensitive to ending radius due to the lack of data points beyond R=16



# Model Comparison

- Our model is stronger than most previous studies
- Similar to Amores et al.
  - More physics
  - Flaring, disk truncation, star formation history, etc.
  - Reassuring check of our model



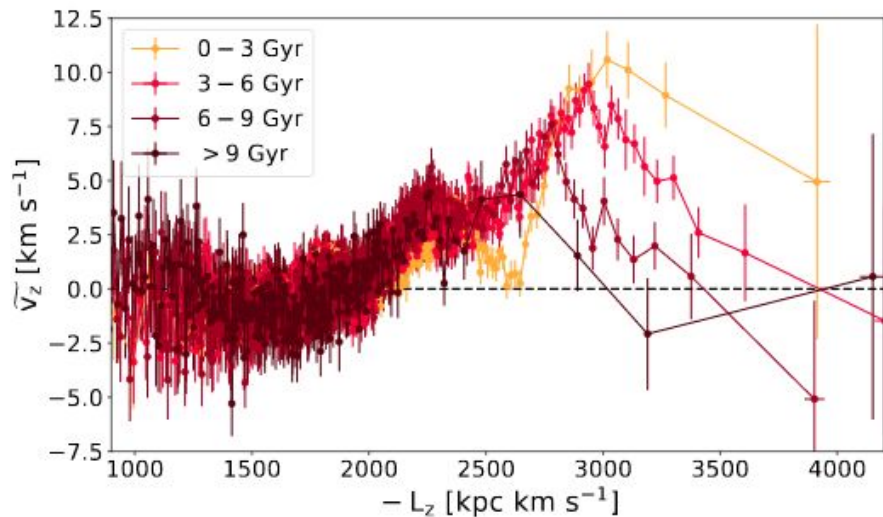
# Model Visualization

Fixed point above the Galaxy

Following the Sun above the Galaxy

# Dynamically Evolving Warp

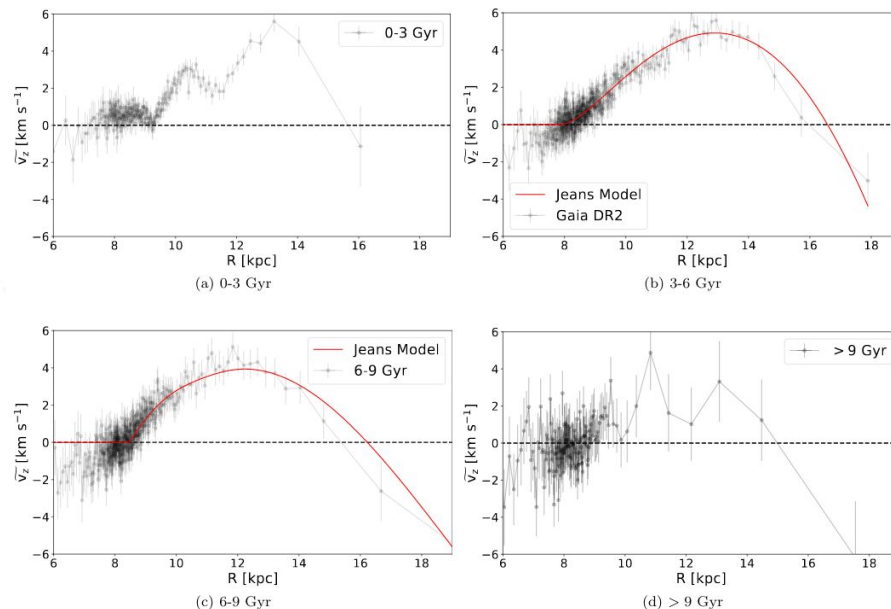
- Age from Sanders & Das (2018)
  - Bayesian isochrone fitting
  - RV, stellar parameters from APOGEE, GALAH, LAMOST, RAVE, and SEGUE
- Warp parameters dependent on age
  - Warp amplitude decrease with age
- Consistent kinematic signature across different age population, except the  $> 9$  Gyr
  - Warp possibly caused by gravitational mechanism



# Dynamically Evolving Warp

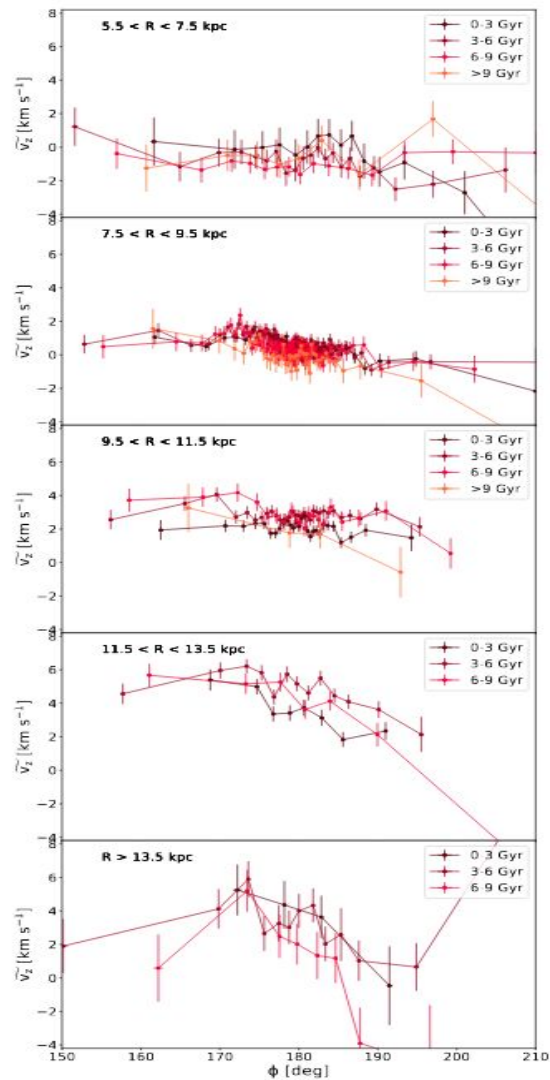
- Precession remains consistent across time

- 0-3 Gyr: Unable to fit
- 3-6 Gyr:  $-11.59^{+0.30}_{-0.25} \text{ km s}^{-1} \text{ kpc}^{-1}$
- 6-9 Gyr:  $-12.19^{+0.49}_{-0.39} \text{ km s}^{-1} \text{ kpc}^{-1}$
- >9 Gyr: dominated by halo stars



# Dynamically Evolving Warp

- Galactic warp remain lopsided across different age population





# Conclusion

- First time discovery of the drop-off in vertical velocity
- Ripples in vertical velocity
- Ripples in radial velocity
- Lopsided Galactic warp
- Consistent modeling of the Galactic warp
- Evolution of warp with time
  - Signature remains consistent for a long time
  - Dynamically evolving
  - Possible gravitational instead of non-gravitational mechanism

Questions?