

Galactic Warp Through the Lenses of Gaia Data Release 2 and the APOGEE Survey

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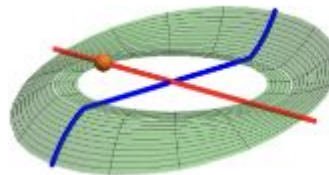
QR Code to ADS



gaia

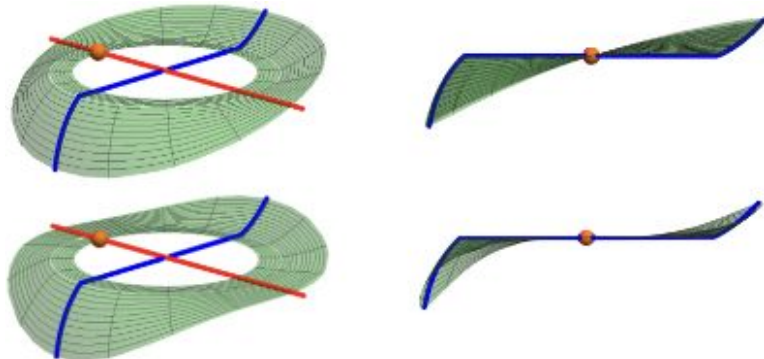


Introduction



- Galactic Warp: the bending of Galactic disk
- Found in majority of spiral galaxies (Bosma 1978)
 - Long-live/repeatedly regenerated
- Origin: under debate
 - Interaction with satellite galaxies
 - Sgr dSph (Ibata & Razoumov 1998; Laporte et al. 2019)
 - LMC, SMC (Weinberg & Blitz 2006; Garavito-Camargo et al. 2019)
 - External torques of dark matter halos (Widrow et al. 2014)
 - Accretion of intergalactic matter (Ostriker & Binney 1989)
 - Misaligned dark matter halo (Sparke & Casertano 1988)
 - Intergalactic magnetic field (Battaner et al. 1990; Gujarro et al. 2010)

Introduction

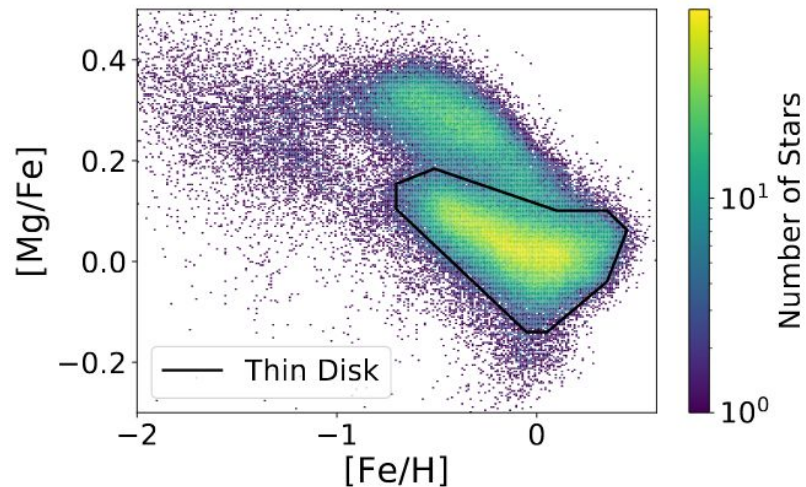


- Geometry: uncertain
 - Shape
 - Sine-lopsided or S-lopsided (Romero-Gomez et al. 2019)
 - Starting radius
 - Inside solar circle
 - J/K-band surface brightness fitting (Drimmel & Spergel 2001)
 - Gaia-TGAS (Schonrich & Dehnen 2018), LAMOST-TGAS (Huang et al. 2018)
 - Outside solar circle
 - Star count analysis (Reyle et al. 2009; Derriere & Robin 2001)
 - Gaia DR2 kinematics (Romero-Gomez et al. 2019)
- This work
 - Kinematics from Gaia DR2
 - Chemistry from SDSS/APOGEE survey (Majewski et al. 2017)
 - Distance from StarHorse (Queiroz et al. 2018)
 - Explore asymmetries in the outer Galactic disk

Data

2 data-sets:

- Gaia
 - Proper motion + StarHorse distance + RV (Gaia)
 - $G < 15$ mag
 - Removal of stars with suspect photometry and RV based on < 4 visits
- Gaia-APOGEE
 - PM + StarHorse distance + RV (APOGEE) + chemistry information
 - Chemically selected thin disk

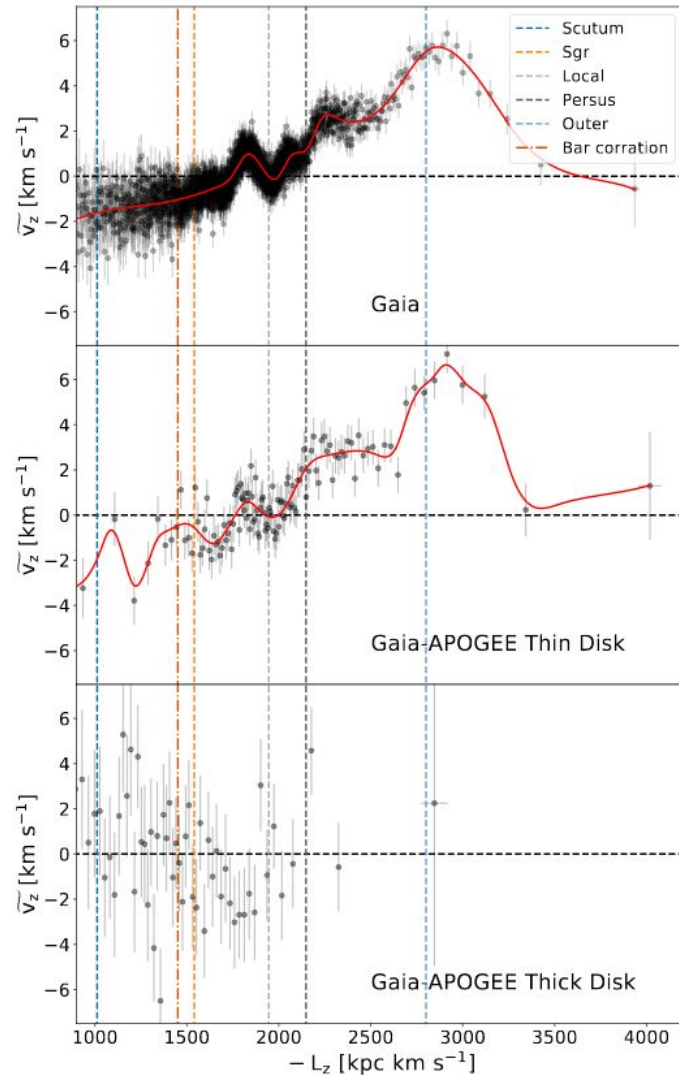


StarHorse distance:

Bayesian estimation of stellar parameters, distances and extinctions with the combination of photometric and parallax information from Gaia, Pan-STARRS1, 2MASS, and AllWISE

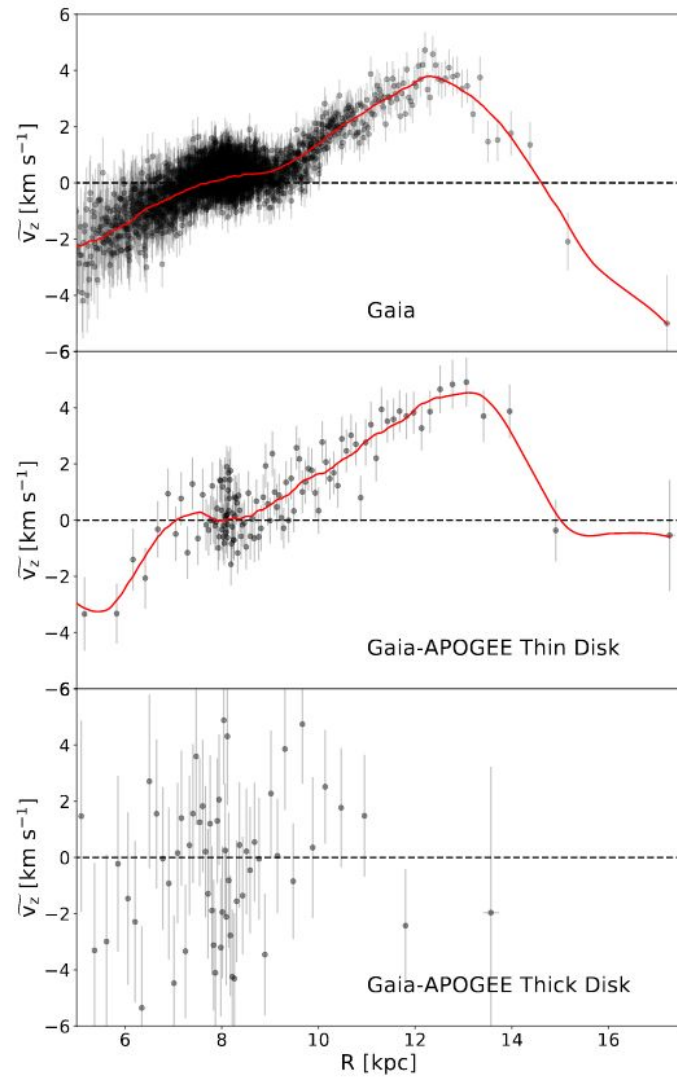
Structures in Vertical Velocity

- V_z increase with L_z
 - $L_z < 2800$ kpc km/s
 - Previously observed and attributed to Galactic warp
- Peak $V_z = 6$ km/s
 - $L_z \sim 2800$ kpc km/s
- Decrease afterwards
- **Discovery of a decline in vertical velocity for the first time**
- Substructure (ripples)
 - $L_z \sim 1700, 2000, 2200$ kpc km/s
 - Some aligned with spiral arms



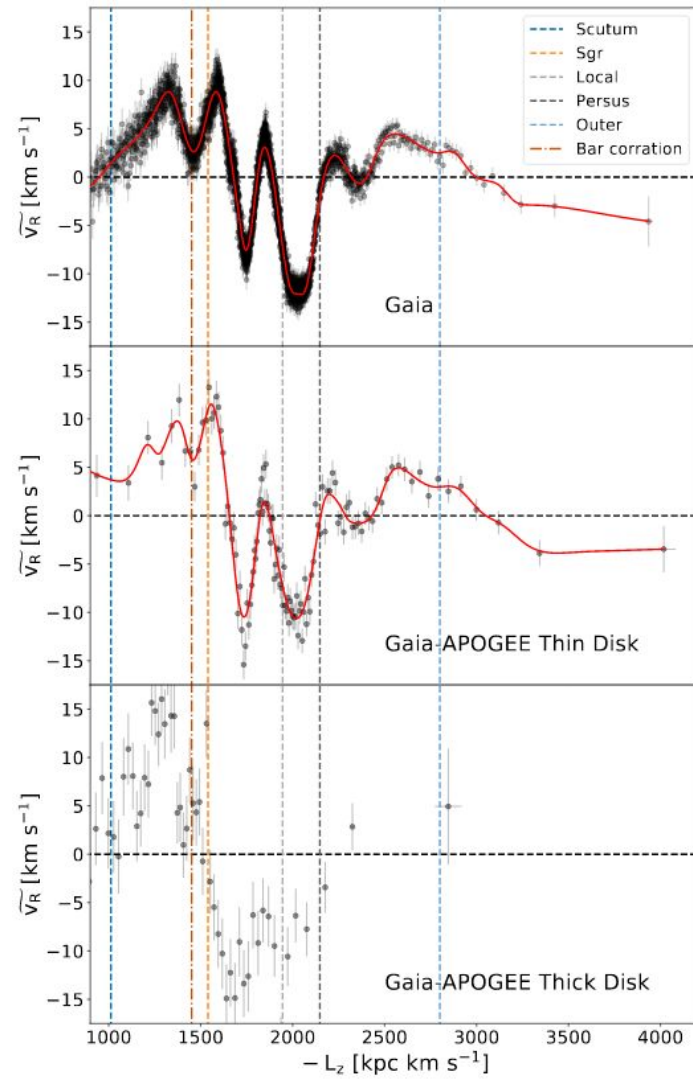
Structures in Vertical Velocity

- Global trend also exists when viewed against Galactocentric radius
- Substructures smeared out since R is not an integral of motion



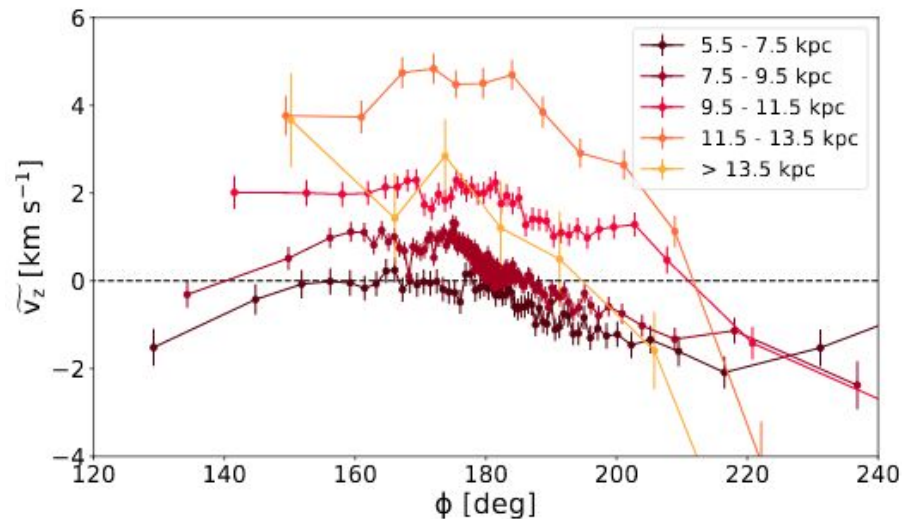
Structures in Radial Velocity

- Large-amplitude (~ 13 km/s) ripples detected
- Phase difference with vertical velocity
- Not all peak/valley aligned with spiral arms



Lopsided Warp

- Romero-Gomez et al. 2019: up and down sides are not symmetric -> lopsided
- Peak velocity at ~ 170 deg
- Rate of change in vertical velocity not symmetric with respect to the peak
- **Lopsided warp through stellar kinematics**



Modeling the Observed Vertical Kinematics

- Non-zero radial velocity
- Warp precessing in galactic-rotation direction
- Axis-symmetry
- Plug into Jeans Equation:

$$\begin{aligned} & \left[\overline{v_z} - \left(\frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos(\phi + \phi_w + \omega_p t) - \right. \\ & \left. \overline{v_R} \frac{dh}{dR} \sin(\phi + \phi_w + \omega_p t) \right] f(R) \frac{dg}{dz} \\ & + \overline{v_R} \frac{df}{dR} g(z - h(R) \sin(\phi + \phi_w + \omega_p t)) + \frac{\partial \overline{v_R}}{\partial R} n' = 0 \end{aligned}$$

- Double exponential law: $n(R, z) = n_0 \exp\left(-\frac{|z|}{z_h} - \frac{R}{R_h}\right)$
- Plug in and crunch: $\frac{\partial \overline{v_R}}{\partial R} = \frac{\overline{v_R}}{R_h} + \frac{\text{sign}[z - z_0]}{z_h} \left[\overline{v_z} - \left(\frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos \theta - \overline{v_R} \frac{dh}{dR} \sin \theta \right]$
- Population is symmetric about z_0 , so average $Z > Z_0$ and $Z < Z_0$

$$\overline{v_z} = \left(\frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos \theta + \overline{v_R} \frac{dh}{dR} \sin \theta$$

Modeling the Observed Vertical Kinematics

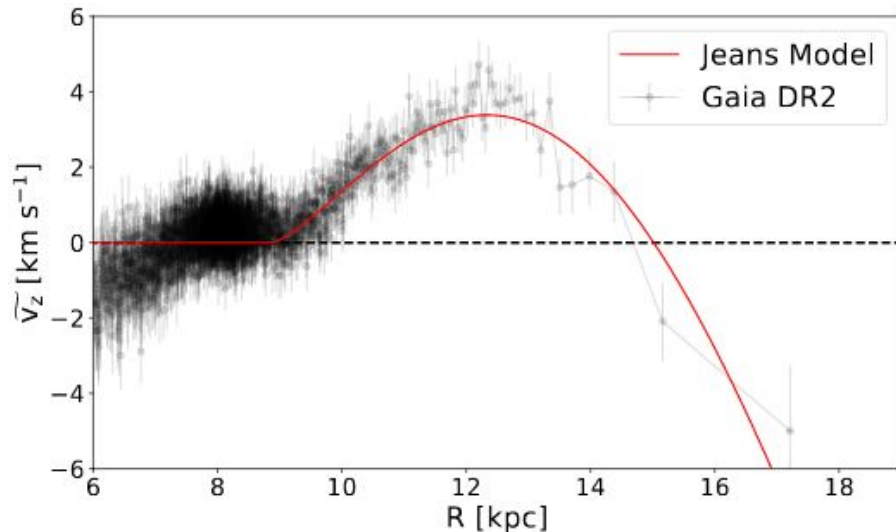
Analytical model:

- Based on Jeans Equation
- Including terms generally ignored

Warp stars at ~ 8.9 kpc

Precession period: ~ 450 Myr

- ~ 2.25 x slower than Galactic rotation at Sun

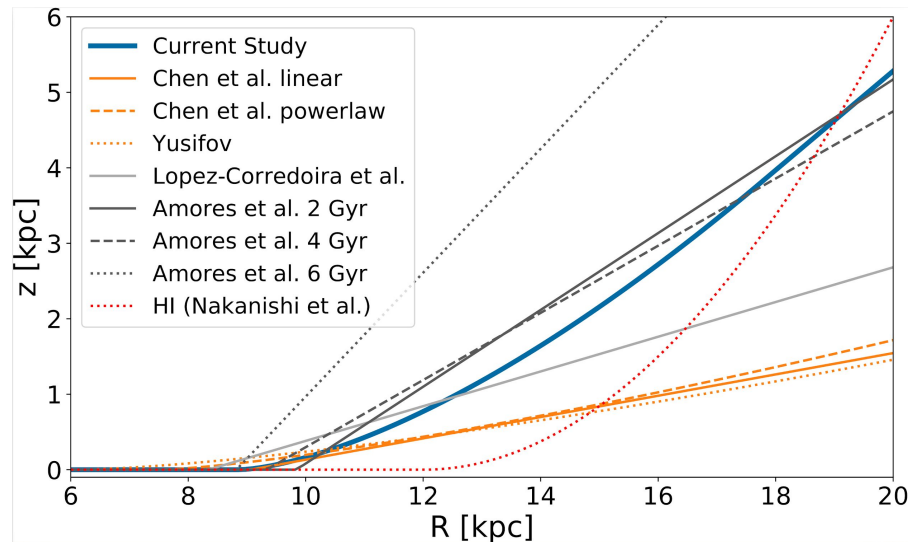


$$\overline{v_z} = \left(\frac{\overline{v_\phi}}{R} + \omega_p \right) h(R) \cos \theta + \overline{v_R} \frac{dh}{dR} \sin \theta$$

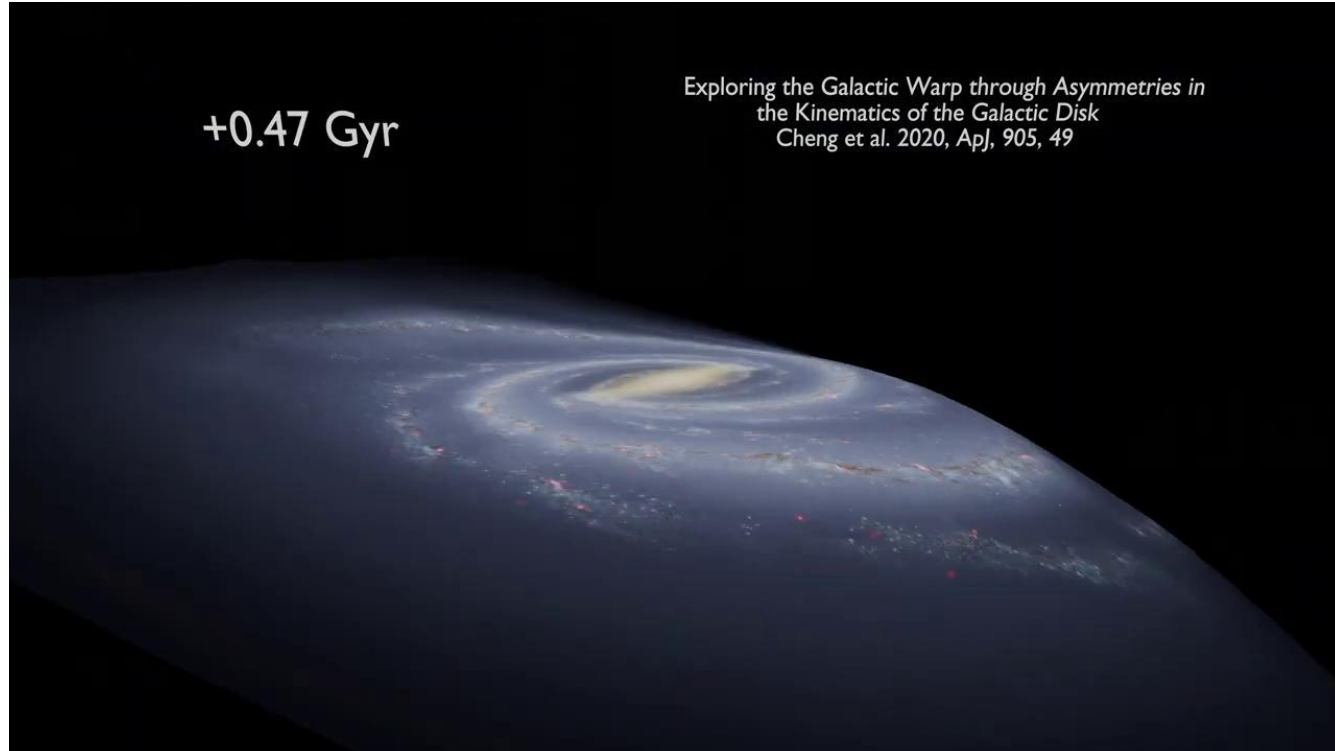
Full derivation in paper

Model Comparison

- Our model is stronger than most previous studies
- Similar to Amores et al. 2017
 - More physics
 - Flaring, disk truncation, star formation history, etc.
 - Reassuring check of our model

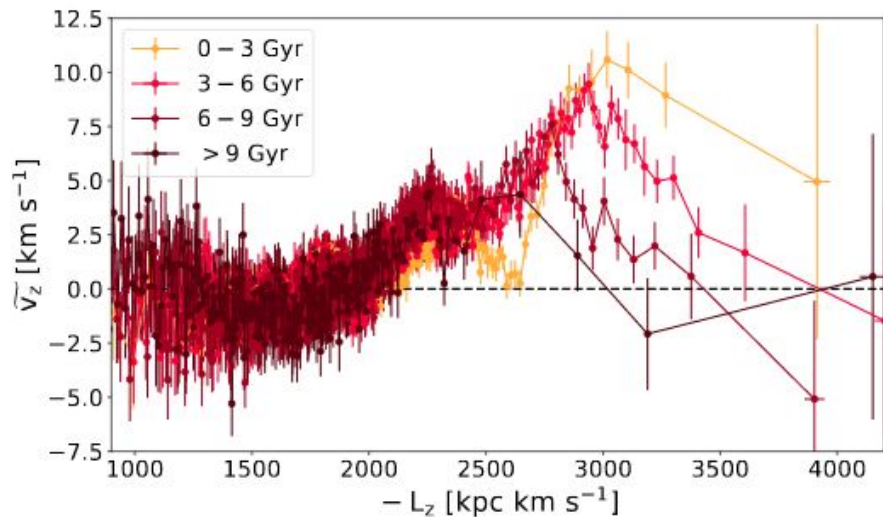


Model Visualization



Dynamically Evolving Warp

- Age from Sanders & Das (2018)
 - Bayesian isochrone fitting
 - RV, stellar parameters from APOGEE, GALAH, LAMOST, RAVE, and SEGUE
- Warp parameters dependent on age
 - Warp amplitude decrease with age
 - Precession rate similar
 - Warp possibly caused by gravitational mechanism within the past 3 Gyr

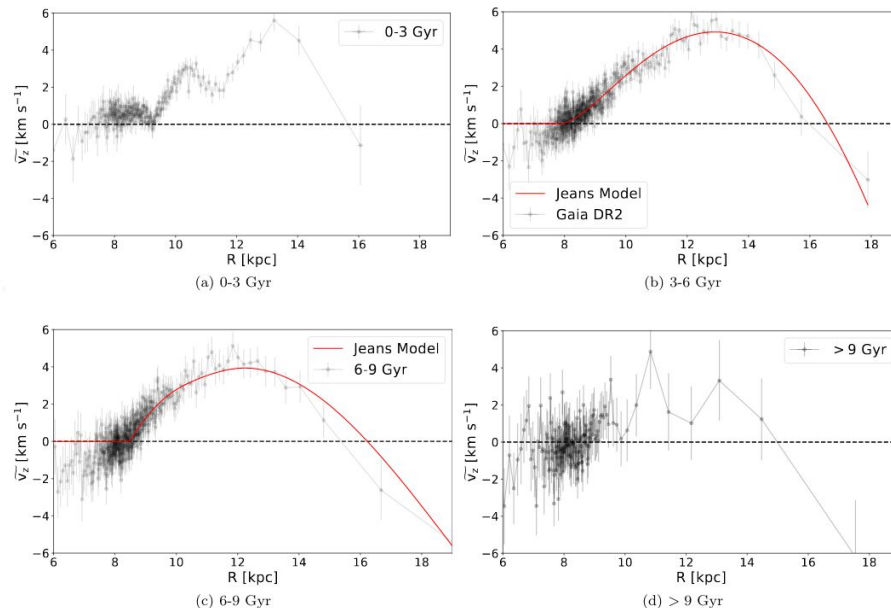


Dynamically Evolving Warp

- Precession remains consistent across time

- 0-3 Gyr: Unable to fit
- 3-6 Gyr: $-11.59^{+0.30}_{-0.25} \text{ km s}^{-1} \text{ kpc}^{-1}$
- 6-9 Gyr: $-12.19^{+0.49}_{-0.39} \text{ km s}^{-1} \text{ kpc}^{-1}$
- >9 Gyr: dominated by halo stars

- Galactic warp remains lopsided across different age population



Conclusion

- First time discovery of a decrease in vertical velocity in the outer part of the galaxy
- Ripples in vertical and radial velocity
- Lopsided Galactic warp
- Able to interpret warp with simple Jean Eqn analytical model
- Evolution of warp with time
 - Younger population, stronger response
 - Possible gravitational instead of non-gravitational origin within the last 3 Gyr

Questions?